

# Flickering of the Vela pulsar during its 2016 glitch

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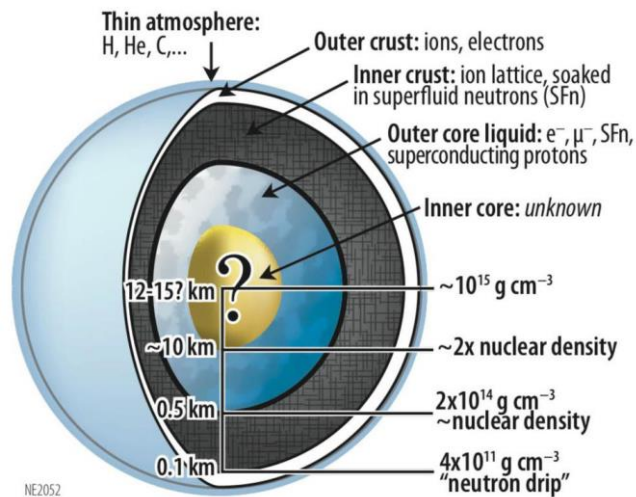
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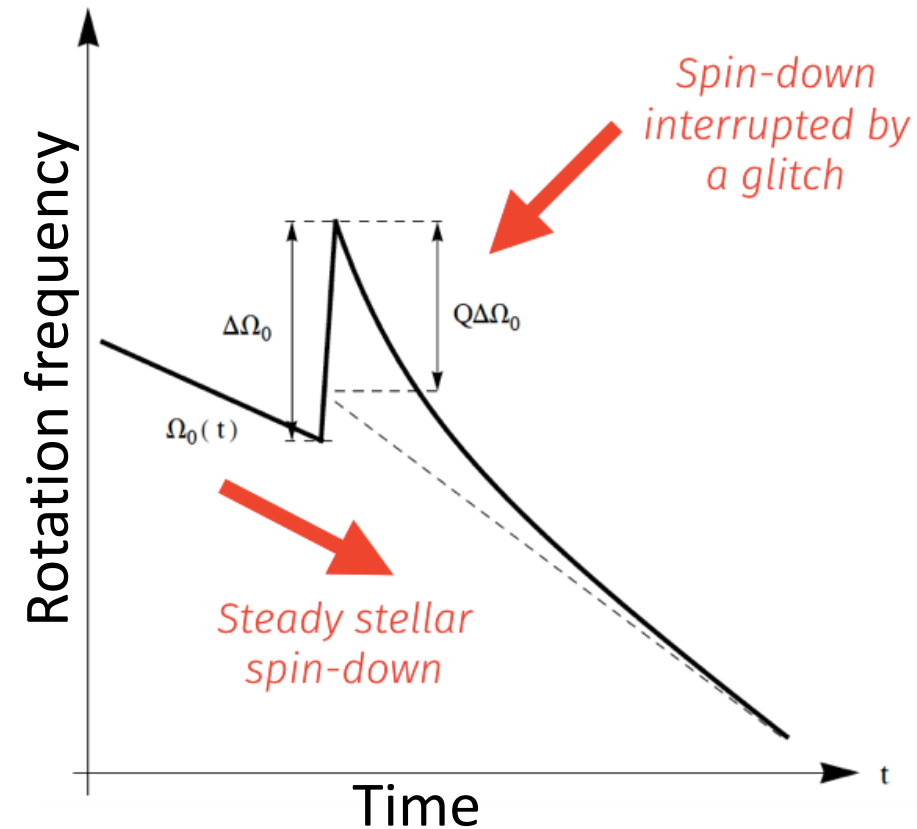


# Radio pulsar glitches

- Sudden spin-up event
- Coupling of the inner and outer crust
- Insights into the interior
- See [Haskell & Melatos \(2015\)](#)

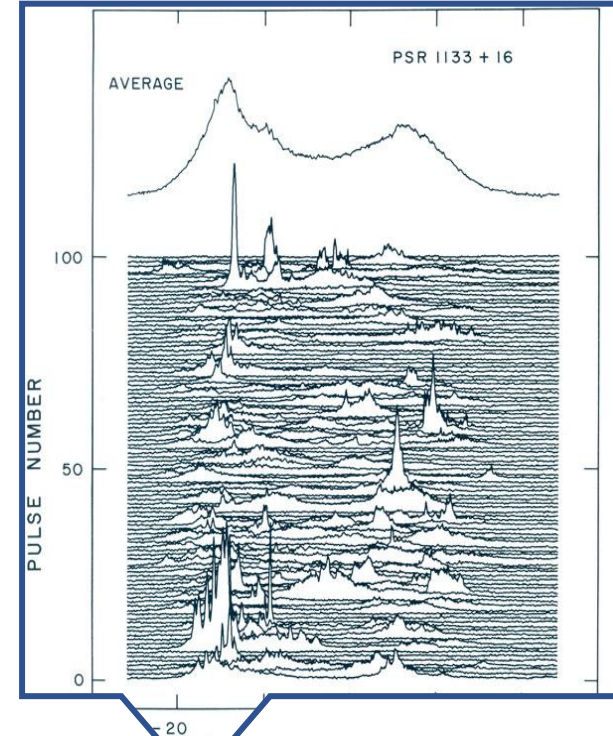
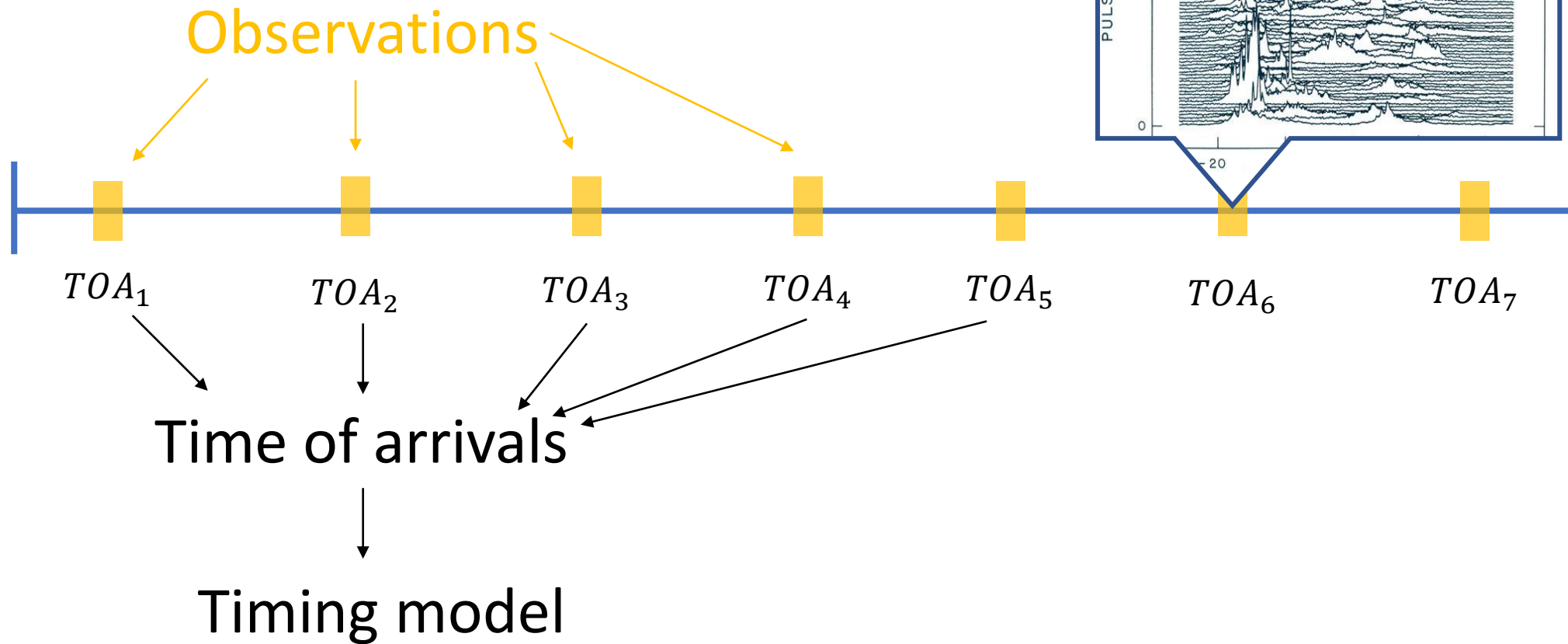


[NASA / GSFC / NICER](#)

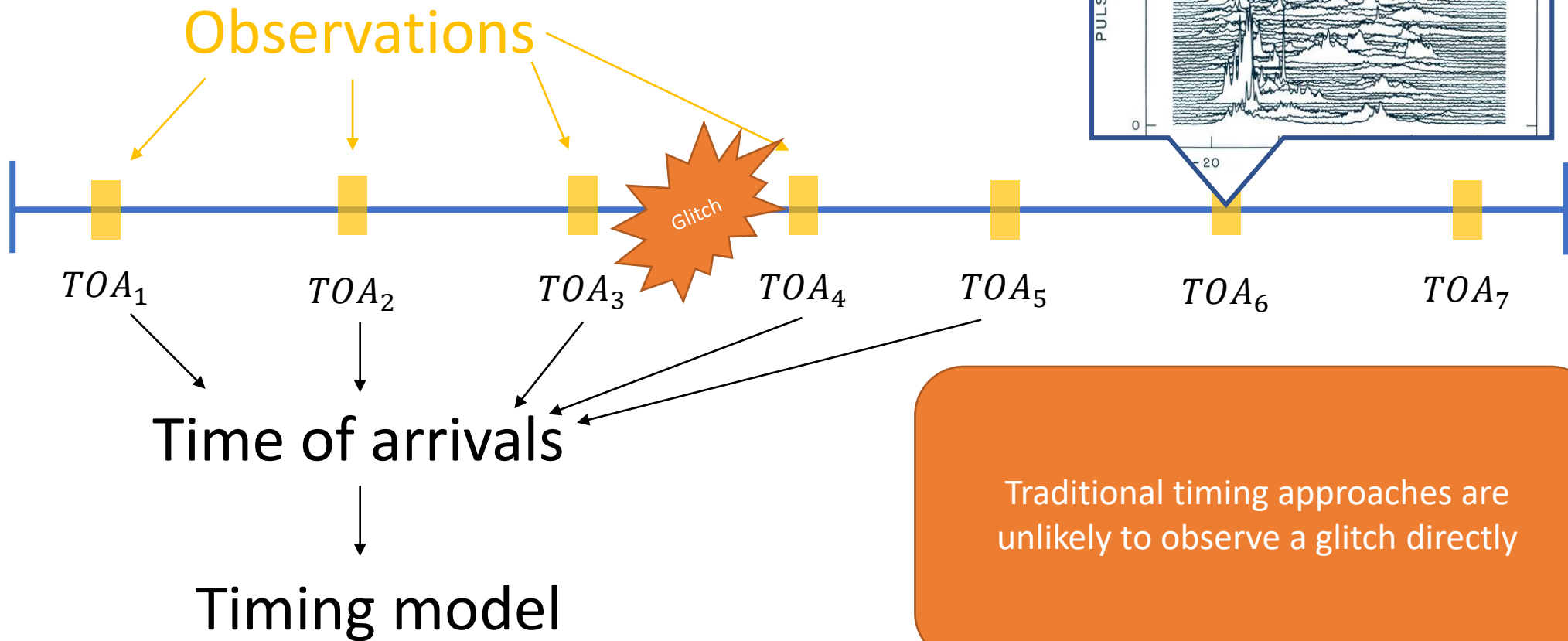


[Middleditch et al. \(2018\)](#)

# Traditional pulsar timing



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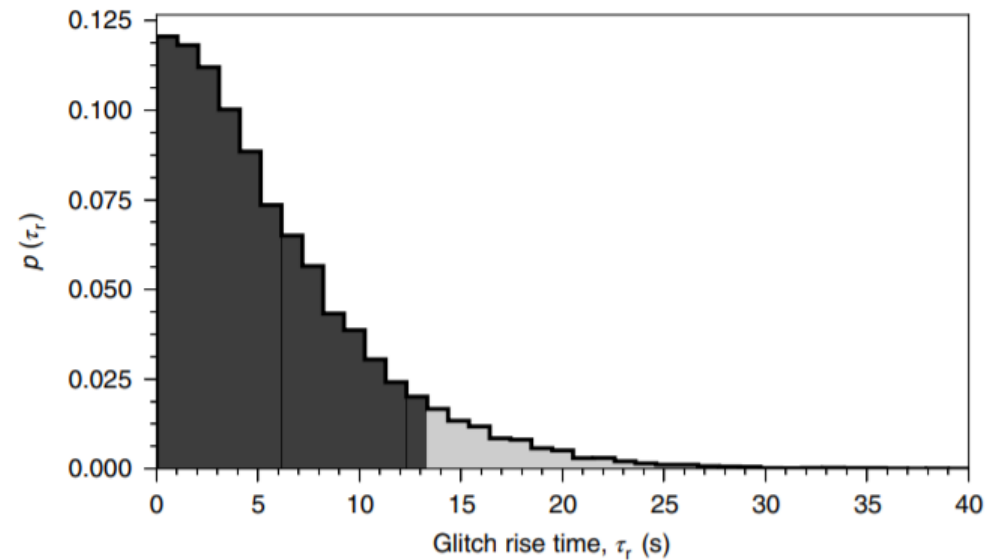
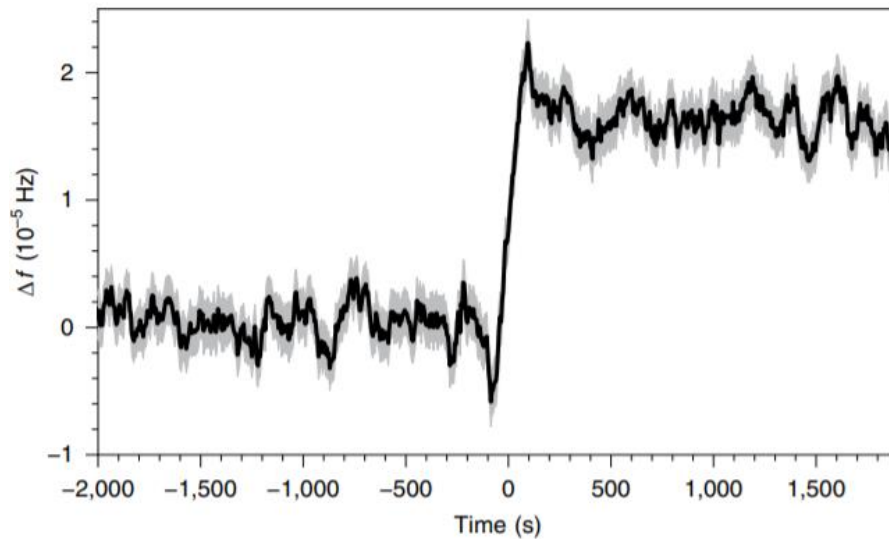
# The 2016 Vela radio-pulsar glitch

- Mt Pleasant Observatory Tasmania
- Constantly surveilling the Vela pulsar
- In 2016 it caught a glitch in real time:  
*“Pulse-to-pulse observations”*
- [Palfreyman et al. \(2018\)](#)



# Seeing the spin-up

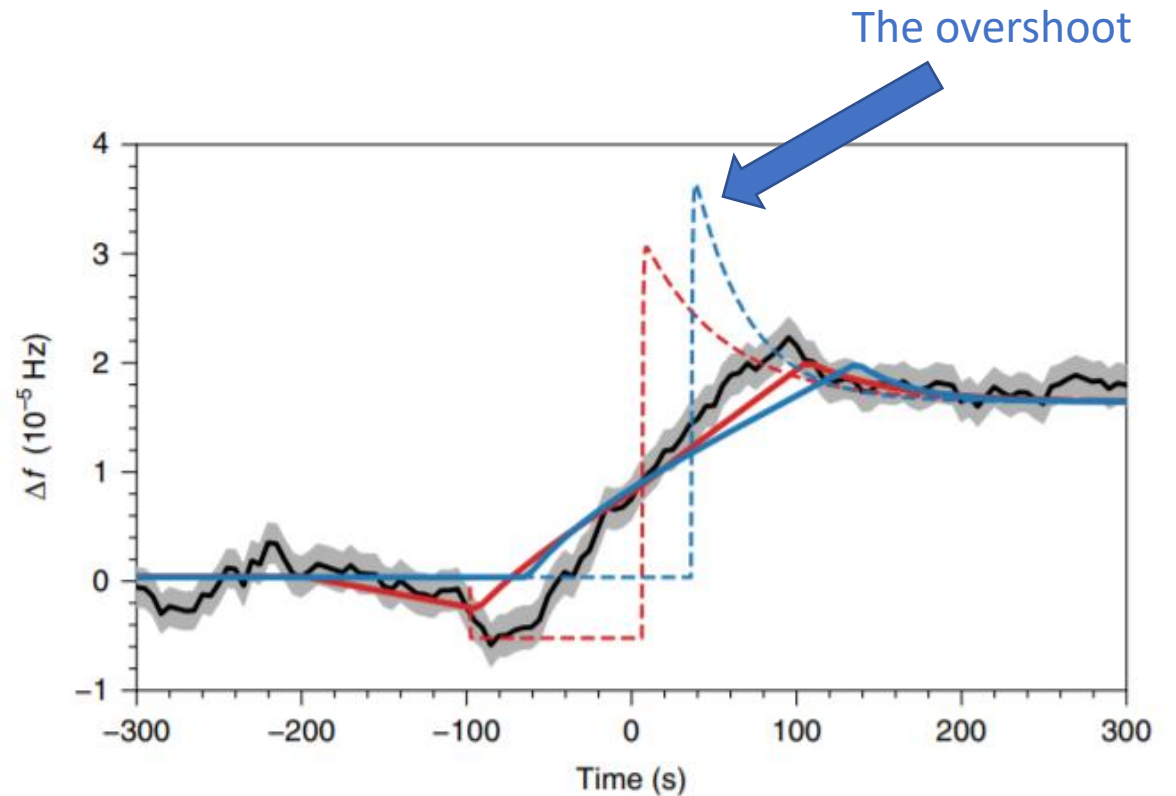
- The data allowed us to see the spin-up itself, for the first time:



[Ashton et al. \(2019\)](#)

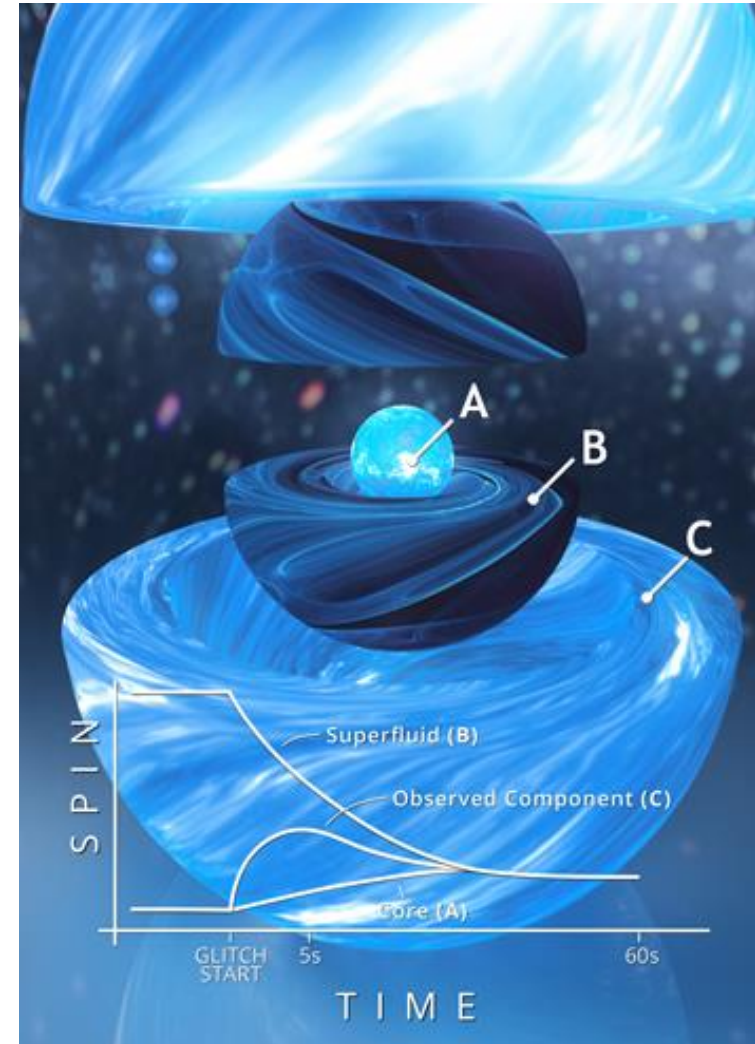
# Using the dynamics to probe the physics

- We fit phenomenological models to infer the glitch properties
- We find overwhelming evidence for an “overshoot”



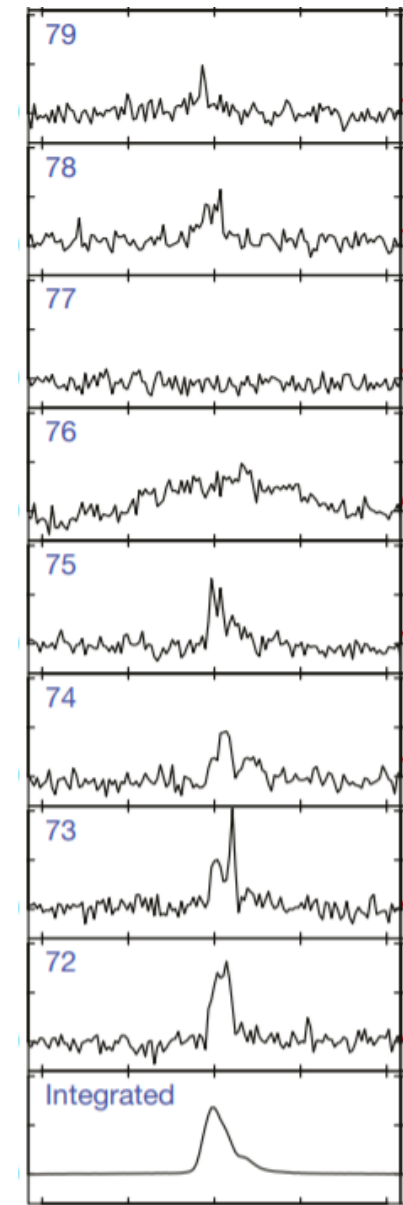
# Evidence for three-components

- The overshoot suggests the existence of three separate components
- This allows direct measurements of the Moment of Inertia and coupling between components:
  - [Pizzochero et al. \(2020\)](#)
  - [Gügercinoğlu et al. \(2020\)](#)
  - [Montoli et al \(2020\)](#)



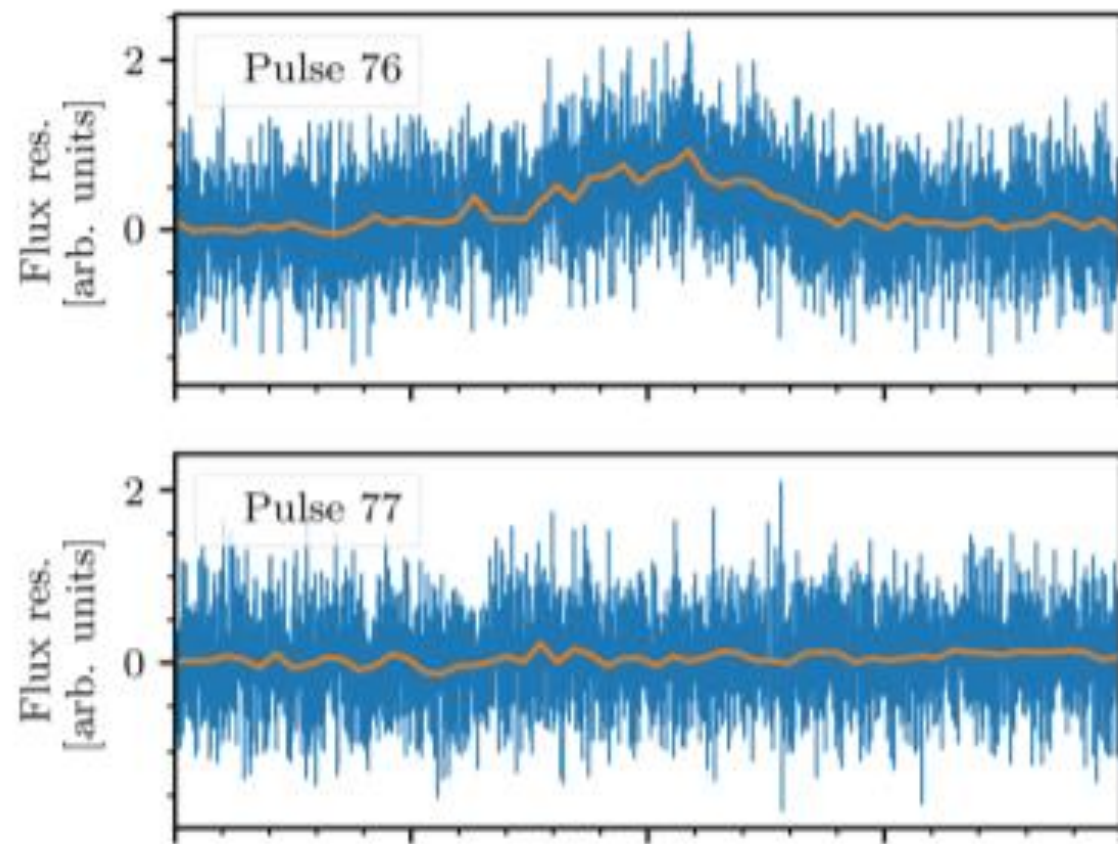
# A null just prior to the glitch

- This data allowed Palfreyman to analyze individual pulses during the glitch.
- While **integrated pulses are stable**: pulsars are known to exhibit significant **jitter**.
- A null, **pulse 77**, occurred just before the glitch.
- This is the first recorded null in the Vela pulsar



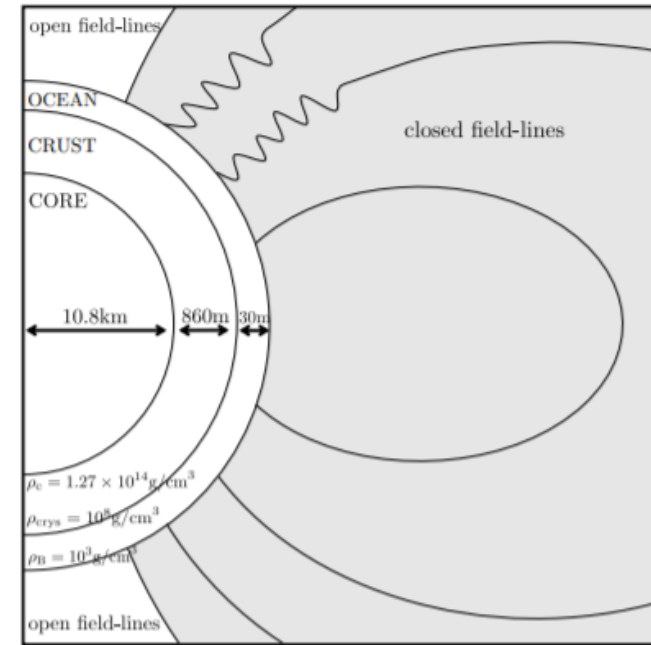
[Palfreyman et al \(2018\)](#)

# The broad pulse and null pulse in more detail

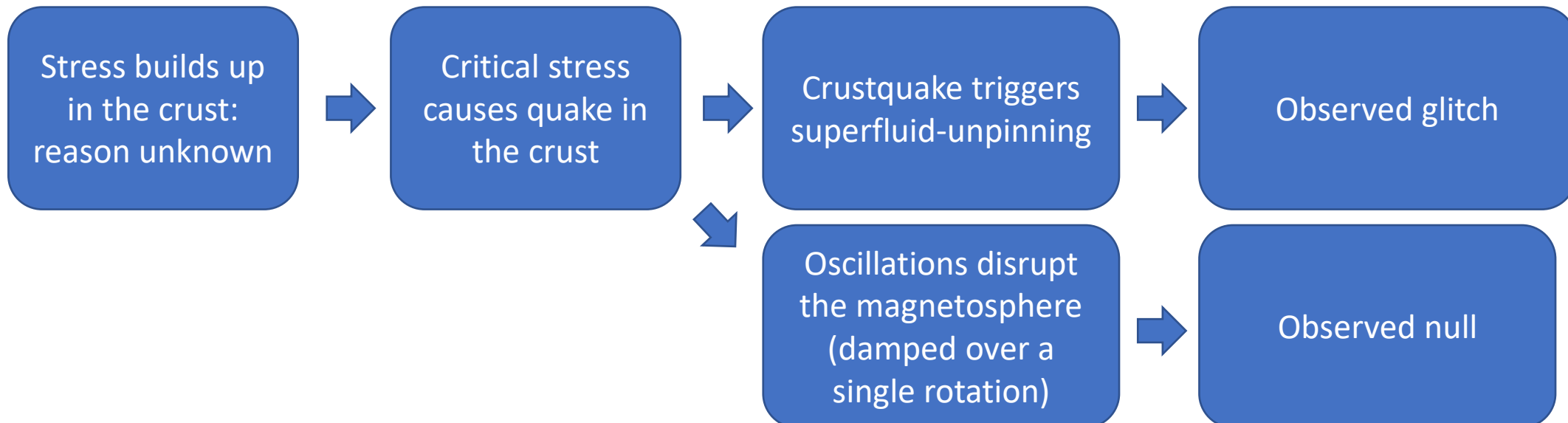


# Implications of the null

- [Bransgrove et al. \(2020\)](#) develop a model connecting the glitch and null:



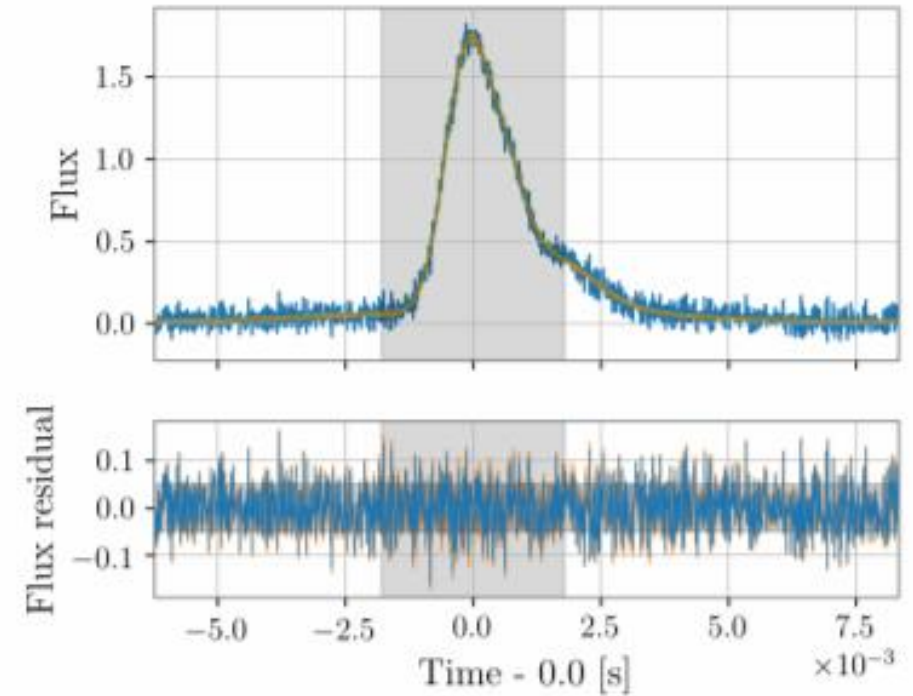
[Bransgrove et al. \(2020\)](#)



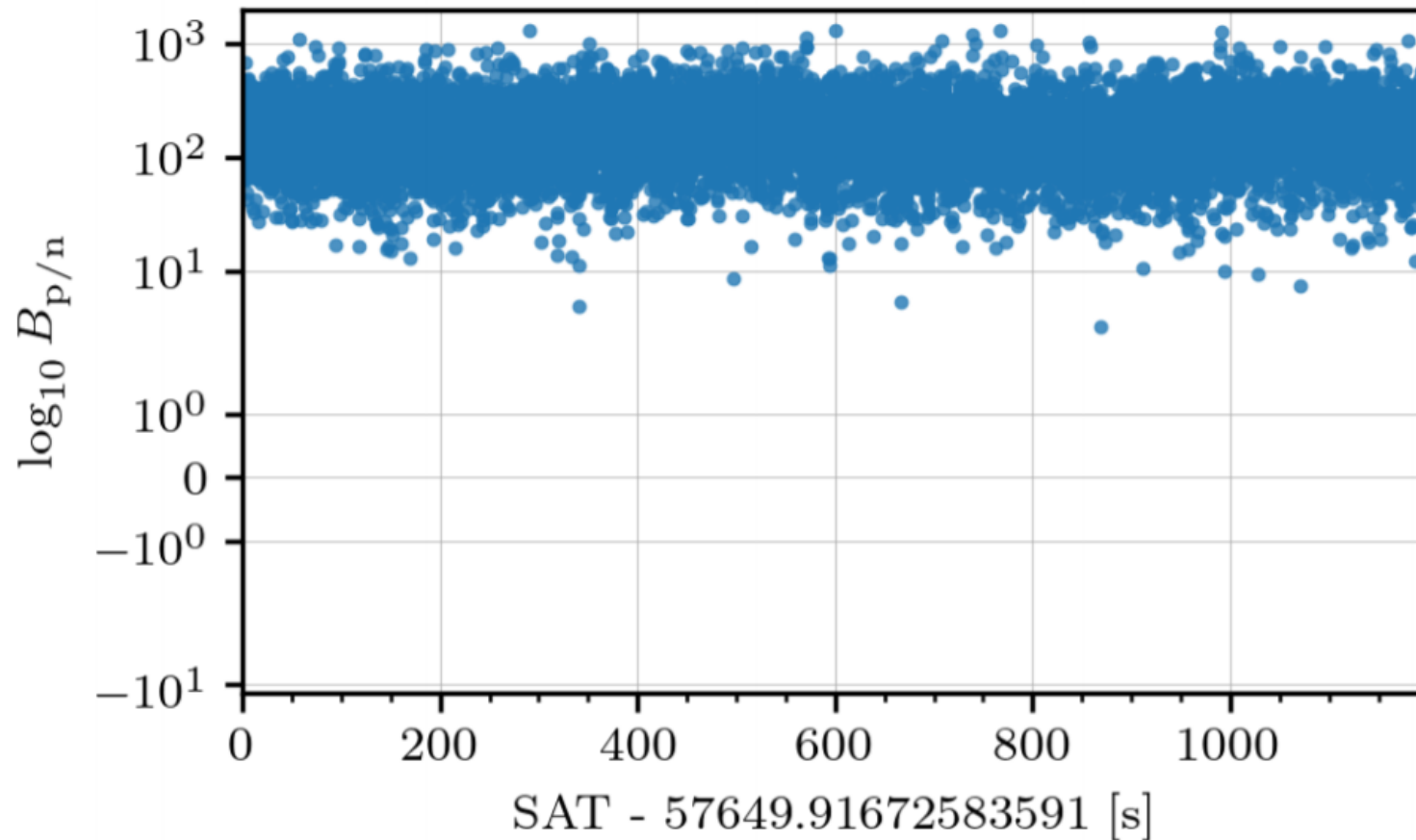
# Reanalyze the raw data

- [Ashton et al. \(2020\)](#)
- We fit phenomenological flux models to the raw data of each individual pulse
- Use the open-source [kookaburra](#) package
- Can calculate a Bayes factor

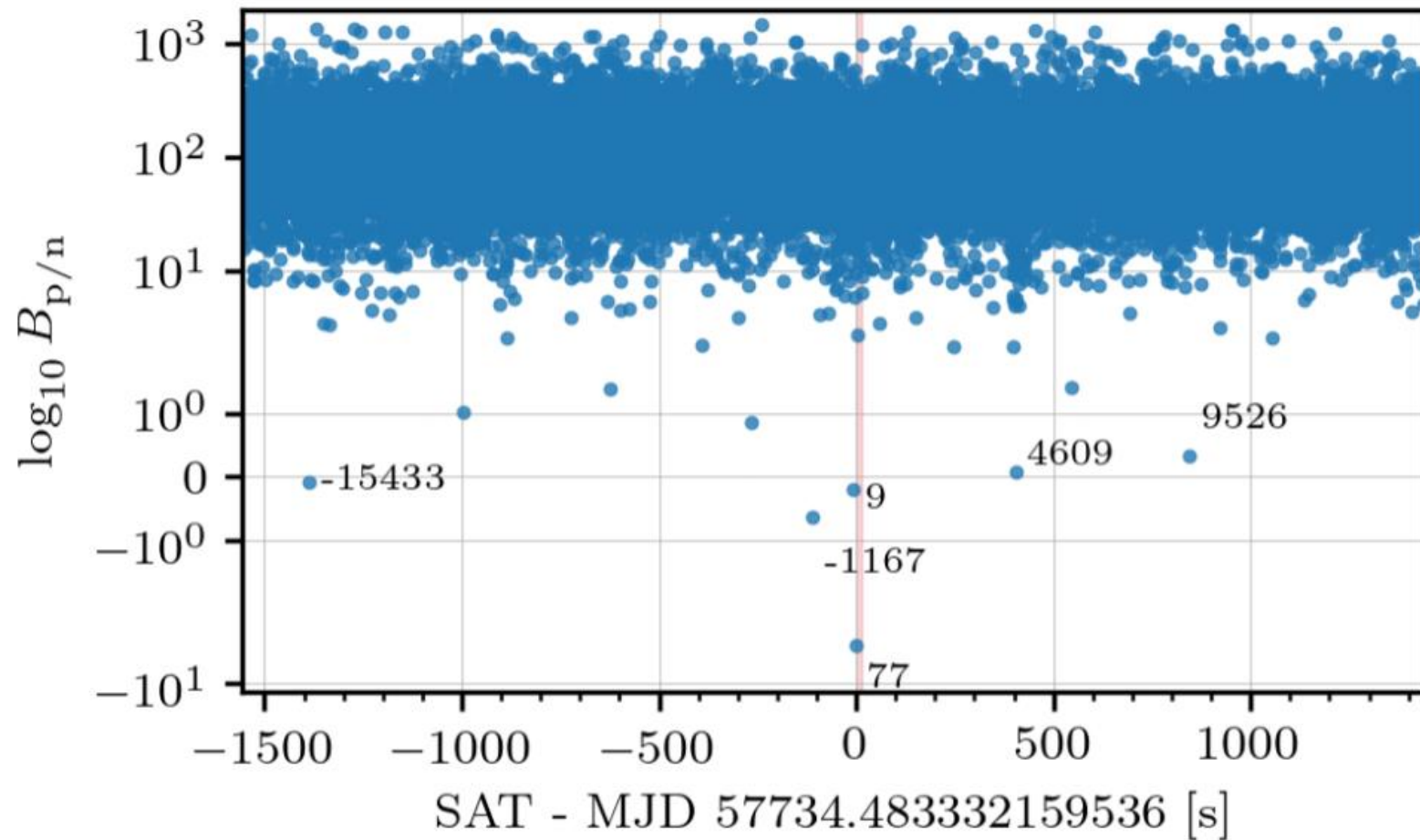
$$B_{p/n} = \frac{P(\text{pulse} \mid \text{data})}{P(\text{null} \mid \text{data})}$$



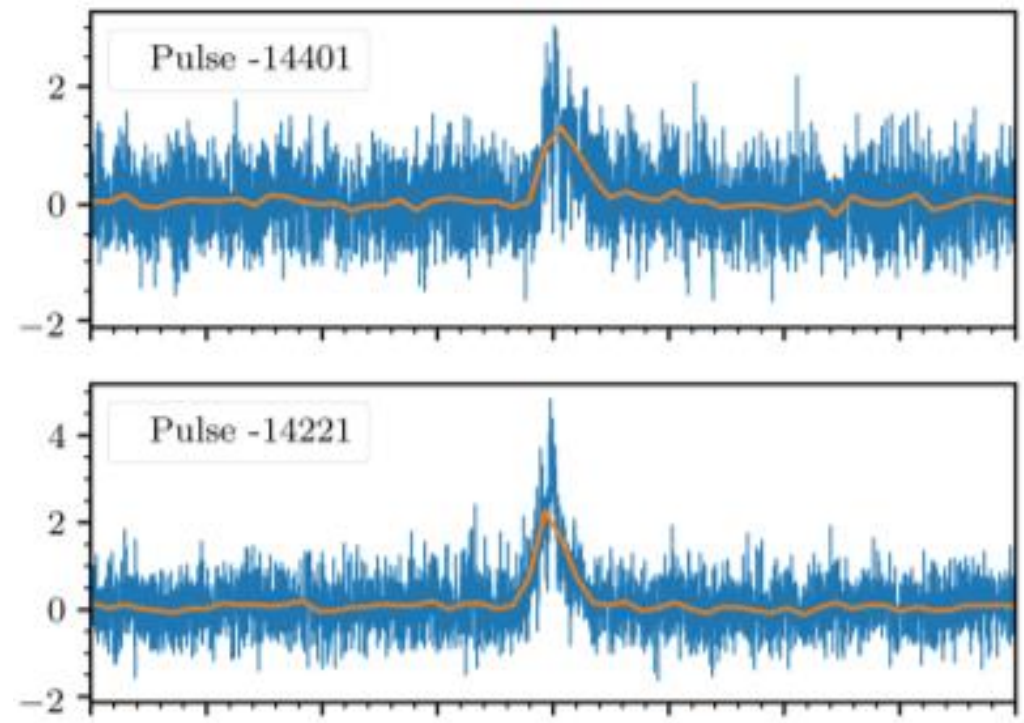
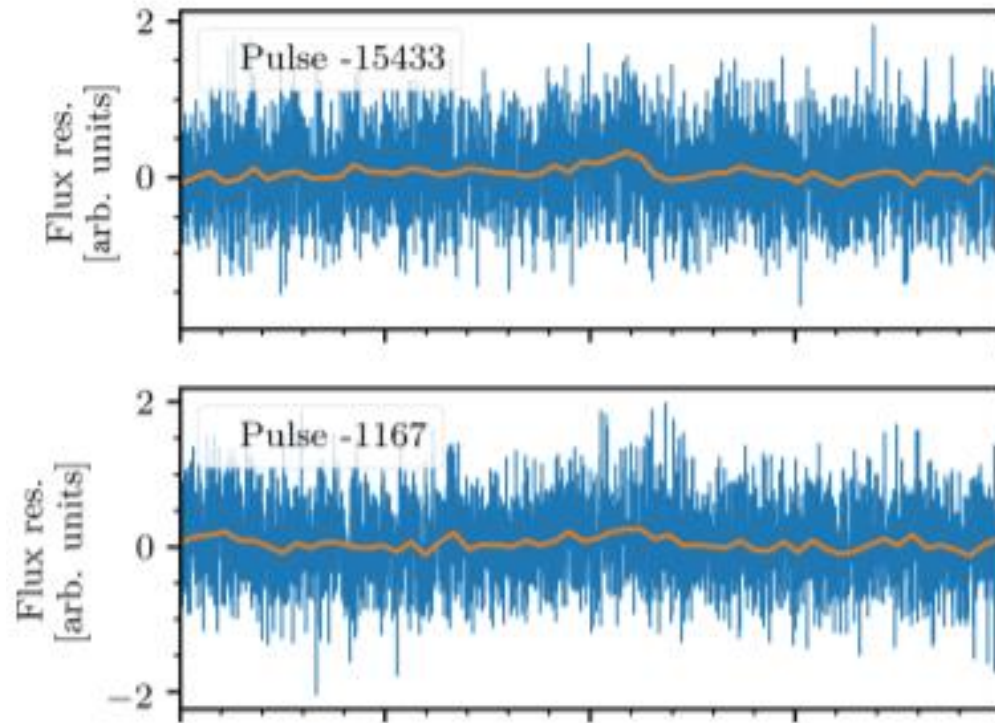
20 minutes of data away from the glitch



# 50 minutes of data around the glitch



# Quasi pulses: what do they look like?



# Conclusions

- Re can re-affirm the existence of the null
- We also find evidence of the existence of “quasi-nulls”
  - What does this mean for the Bransgrove et al interpretation?
  - Could the null + quasi-nulls instead be explained by a “magnetospheric storm”
- Currently analyzing a larger quantity of data
- Extended activity could be used as an early-warning system